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Laser-based satellite positioning system for space gravitational wave detector DECIGO



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ABSTRACT

We propose the inter-satellite laser positioning system for precision formation flying. Our system consists of a single frequency laser and acousto-optic deflectors, and the remote spacecraft can obtain their angle information by frequency counting of the laser beat signal. The accuracy of the angle measurement is evaluated to be 0.004 mrad, and we expand this system to two-dimensional angle measurement. The distance between S/Cs can be also measured by using microwave phase in the detected beam, and length resolution of 6.25 m is obtained from our preliminary test. We plan to apply this system to the initial alignment of three S/Cs for the precision formation flying of Japanese space gravitational wave detector DECIGO project.

Keywords: laser ranging, formation flying, angle measurement, gravitational wave detector

1. INTRODUCTION

Precision formation flying of the space crafts (S/C) is one of the essential techniques for the future space science missions such as very long-baseline infrared interferometers [1][2], a satellite gravimetry mission [3], long-focal-length X-ray or gamma ray telescopes [4][5][6] and much larger-size satellite network for testing general relativity [7][8]. In these mission, effective aperture size or focal length of the telescope is dramatically expanded to improve angle or image resolution. Among these formation flying missions, space gravitational wave detectors would have highest degree of technical difficulties from their much longer formation size and precision.

Gravitational wave had been predicted in 1916 by A.Einstein in his general theory of relativity which would be generated by the acceleration of mass such as explosion of super novae or inspiral and coalescence of blackhole binaries. Due to its extremely small strain of lower than $\delta l/l < 10^{-23}$, direct detection of gravitational wave is very difficult. In 2016, the first direct detection of gravitational wave had been achieved by advanced LIGO [9], and successive detections have been reported from ground-based gravitational wave detectors. The ground-based gravitational wave detector is a Fabry-Perot Michelson laser interferometer with the arm length of 3-4 km, and has strain sensitivity of 10^{-23} level around 10 to 1 kHz. Since the detection band of the ground-based gravitational wave detector at lower frequency range is limited by its finite arm length and seismic noise from the ground, space gravitational wave detectors are planned to detect lower frequency gravitational wave which have much longer armlength and are operating in seismic noise free condition. Europe and US space gravitational detector named LISA consists of three S/C each of which separate more than 2,500,000 km and are triangle-shaped constellation flying [10]. In LISA the distortion of space is measured by using laser transponder for detecting gravitational wave at lower frequency around 1 mHz. In Japan, we promote another space gravitational mission named DECIGO (DECi-hertz Interferometer Gravitational wave Observatory) consists of three S/Cs forming equilateral triangle-shaped laser interferometer with the arm length of 1000 km, whose conceptual design is shown in Fig.1 [11]. DECIGO contains three differential 1000-km Fabry-Perot laser interferometer with the arm length of 1000 km and the including angle of 60 degree in its triangle-shape formation, and strain sensitivity of $dL/L < 10^{-23}$ at the frequency range between 0.1 to 1 Hz for detecting gravitational wave from the inspiral of intermediate-mass blackhole binaries or primordial background gravitational wave from the early universe.

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DECIGO is planned to launch in mid 2030s. Before launching DECIGO, we plan another milestone mission named B-DECIGO, which is down-sized mission of DECIGO with the arm length of 100 km [12].

In order to obtain such high strain sensitivities, the length of each interferometer arm should be identical with the length difference of lower than 0.1 % for improving common-mode noise reduction (CMNR) because higher CMNR reduces the influence of the intensity and frequency noise of the light source on the strain sensitivity of the interferometer. For initial alignment of three S/Cs after releasing them from the launching vehicle and keeping their relative position in the observation, we proposed novel satellite positioning system based on the combination of a single-frequency laser, acousto-optic modulator and frequency counting. In the previous conference (ICSO2018), the basic idea of our angle measurement system was introduced [13]. Our angle measurement system has been expanded to add two-dimensional angle measurement and length measurement functions, the details of which are presented in the current document.

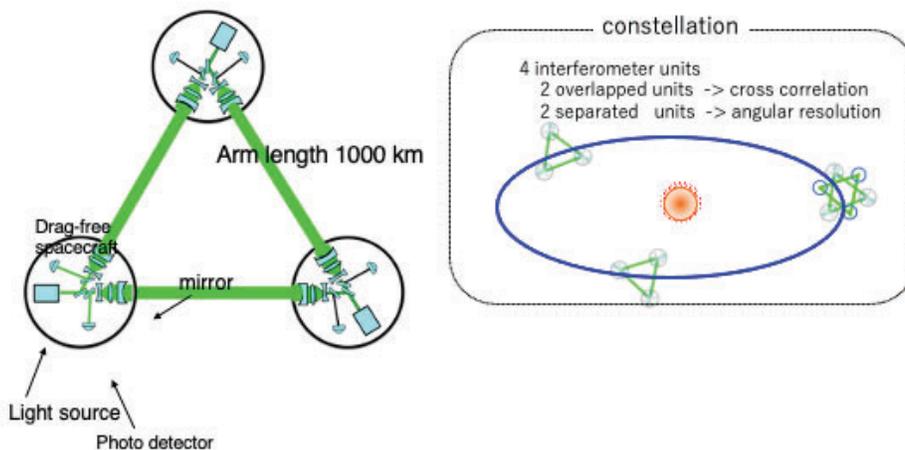


Figure 1 Conceptual design of DECIGO. In the final stage four units of DECIGO will be launched around-synchronous orbits for improving angle resolution and sensitivity

2. ANGLE MEASUREMENT

2.1 principle

Acousto-optic deflector (AOD) is the optical diffraction device which utilizes the interference between acoustic wave and light. When microwave signal is applied to AOD crystal, piezo-electric transducer attached to AOD crystal generates compression wave inside crystal to make Bragg grating. The laser beam introduced into AOD generates 1st diffraction beam whose diffraction angle, θ , is proportional to the driving frequency of AOM, f_m as

$$\theta_m = (\lambda / V_a) f_m$$

V_a and λ in this equation are the acoustic velocity in the crystal and the wavelength of the diffracted laser, respectively. Since the frequency of the 1st diffracted beam is shifted as f_m due to Doppler effect, the frequency shift of the diffracted beam is proportional to the diffraction angle.

The leader S/C emits the first diffracted beam along the reference line on which the follower S/C should be located, and the diffraction angle is scanned by changing driving frequency, f_m . When the follower S/C detects scanned diffracted beam, it obtains the angle information from the frequency of detected diffraction beam. The frequency shift is obtained as the beat note frequency in the diffracted beam

2.2 Experimental set-up

Schematic diagram of the angle measurement system is shown in Fig.2. Laser light with the wavelength of 515 nm is introduced into a polarization beam splitter (PBS1) to divide into signal and local beam. The power ratio of the signal and the local beam is controlled by a half-wave plate inserted before PBS1. The signal beam is introduced into AOD whose 1st diffracted beam is selected by a spatial filter and is propagated toward the remote site. The local beam from PBS1 goes into a local arm1 where acousto-optic modulator (AOM1) is located. AOD1 and AOM1 are driven by the same frequency, f_m , generated by a computer-controlled direct digital synthesizer (DDS). The 1st diffracted local beam from AOM1 is reflected by a concave mirror back into AOM1 again to generate 1st diffraction beam. By optimizing beam mode and alignment, this double-pass configuration makes $2f_m$ -frequency-shifted local beam which goes back to the PBS1 along the same line as of the input local beam. A quarter-wave plate (QWP1) rotates the polarization of the local beam from P to S so that the returned local beam is reflect by PBS1 into the local arm2. In the local arm 2, the local beam is reflected back into PBS1 to collinearly overlap with the signal beam by using a combination of QWP2 and mirror, and is introduced into AOD1 to make diffraction beam with the signal beam. Therefore, the remote site detects both diffracted signal beam with frequency of $f_s=f_0+f_m$ and the local beam with the frequency of $f_l=f_0+3f_m$ to obtain $2f_m$ beat signal between them.

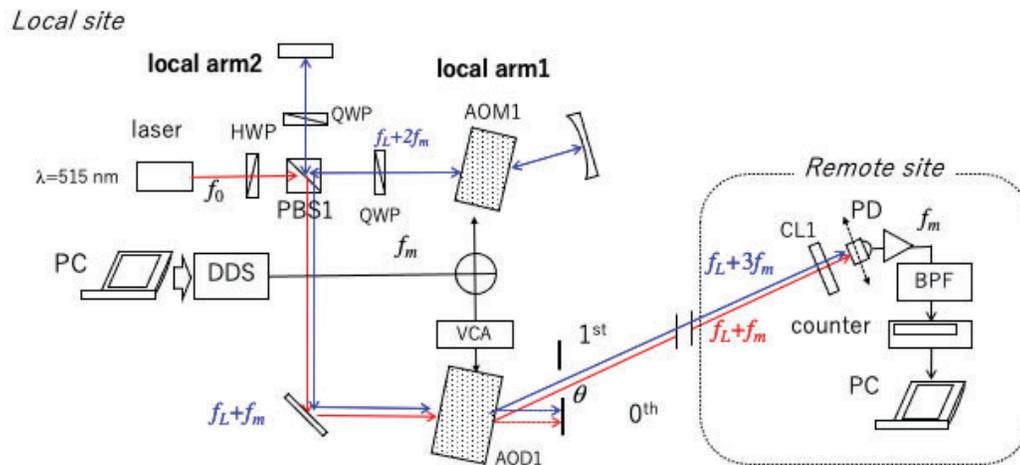


Figure 2 Experimental set-up of the angle measurement system. Red and blue arrows indicate signal and local beam, reflectively. At the remote end, photo detector detects beat note between local and signal beam. DDS: direct signal synthesizer, VCA: voltage-controlled oscillator, BPF: band-pass filter.

At the remote site, the diffracted beam is detected by a fast photo detector for counting beat note frequency and also dc detected power. As the diffraction efficiency of the AOD varies with the driving frequency, the power stabilization servo is applied to AOD1. A part of the diffracted light is pick off, and the power fluctuations of the diffracted beam is suppressed by controlling applied microwave power to AOD. In activating negative feedback servo loop, the diffracted beam power is kept constant while the beam angle is scanned.

2.3 Two-dimensional scanning

For searching the follower satellite, two-dimensional beam scanning is necessary. Therefore, another acousto-optic deflector (AOD2) is added. AOD2 is set just after AOD1 whose crystal axis is orthogonal to that of AOD1, and is driven by f_n with another DDS. The beam propagation can be controlled to arbitrary direction on the 2-D plane by setting f_m and f_n independently. However, the remote S/C should get two information for determine his position on the 2-D plane (angle information of x and y axis). In order to carry two frequency information, f_m and f_n , on one diffracted beam, another acousto-optic modulator AOM2 is added in the local arm 2 which is driven with the frequency of f_n . The local beam from

the local arm 1 is introduced into AOM2, and both 0th and 1st diffracted beam is selected by a spetal filter, and they reflect back into the AOM2 again and only one beam is selected by a spatial filter which go back collinearly with the input beam (0th-to 0th and 1st to 1st diffracted beam). In the consequence, there are three frequency components staying in the diffracted beam from AOD1, $f_0+f_m+f_n$, $f_0+3f_m+f_n$ and $f_0+3f_m+3f_n$. At the remote site, three beat notes with the frequency of $2f_m$, $2f_n$ and f_m+2f_n are obtained. Therefore, the remote site can know their 2-D angle information from $2f_m$, $2f_n$ beat frequencies by detecting one scanned beam.

2.3 result and discussion

For the preliminary test, we set a three-fold 10-m optical path between the local and remote site, and the one-dimensional angle resolution of this system is evaluated. At the local site, the driving frequency of AOD1, f_m , is swept for 5 MHz around 75 MHz to scan diffraction beam angle. The diffracted beam is detected at the remote site, 10.35 m apert from the local site, and the detection system moves for 2 mm along the scanning line on the vertical plane to the optical axis. The detection system has an aperture with the diameter of 1/2" and the transmitted light through the aperture is focused on a photo detector (Hamamatsu S5978) with a 50-mm focal length lens. The detection system obtains both the beat note frequency and DC detected power continuously during beam scanning, and the center frequency of the beat signal is determined at the peak DC voltage. The frequency data are acquired in every 0.5 mm steps. From the 5 times averaging data, the angle precision of 0.0035 mrad with the angle resolution of 0.017 mrad are obtained. The angle resolution limit in this measurement comes from the pointing fluctuation of the diffracted beam which is caused from the vibration of the mirrors and air turbulences. The angle resolution would be improved by using direct path under the vacuum conditions. Theoretical angle resolution limit is determined by the resolution of the frequency count, which comes from the SNR of the beat signal. In order to obtain the angle resolution of 0.001 mrad, the frequency resolution of 10 kHz is required. For satisfying this resolution limit, more than $5\mu\text{W}$ of the detected power is necessary. In the case of DECIGO, beam diameter from the leader satellite is expanded to 28.5 m at the follower satellite due to diffraction, and 10 mW of the beam power is calculated to be high enough for obtaining angle resolution of 0.001 mrad.

3. LENGTH MEASUREMENT

For acquiring or keeping precision formation flying in DECIGO project, not only angle but also distance between spacecrafts should be measured. Currently, time-of-flight method (TOF) is widely used for the length measurement in space [14], and recent progress of optical frequency comb improves the resolution of long-distance measurement [15]. In these methods, ultra-short pulse light source and precision timing or phase measurement are required. We propose length measurement method which can be installed in our angle measurement system. Let me introduce the principle of our length measurement method. When the electric-magnetic (EM) wave with the wavelength of λ is travelling from the local to the remote site, the distance between two sites, L , can be written as

$$L = \left(N + \frac{\phi}{2\pi}\right)\lambda = \left(N + \frac{\phi}{2\pi}\right)\frac{c}{f}$$

where c , f , N and ϕ are speed of light, frequency of EM wave, number of waves in L and residual EM phase, respectively. When laser frequency is changed from f_1 to f_2 , this equation is written as

$$L = \left(N_1 + \frac{\phi_1}{2\pi}\right)\frac{c}{f_1} = \left(N_2 + \frac{\phi_2}{2\pi}\right)\frac{c}{f_2}$$

$$f_1 - f_2 = \left\{ \left(N_1 - N_2\right) + \frac{\left(\phi_1 - \phi_2\right)}{2\pi} \right\} \frac{c}{L}$$

Therefore, the length L can be obtained from the number difference ΔN and phase difference $\Delta\phi$ as

$$L = \frac{(N_1 - N_2) + \frac{(\phi_1 - \phi_2)}{2\pi}}{f_1 - f} c = \frac{\Delta N + \Delta\phi / 2\pi}{\Delta f} c \quad (1)$$

When the frequency varies continuously for Δf , ΔN and $\Delta\phi$ can be measured from the variation of the phase change at the remote end.

According to this principle, the distance between two sites can be measured as follows

The EM wave is travelling toward the remote site, and at the remote site, the detected EM wave is mixed with local oscillator to demodulate for obtaining the phase at the remote end. As the frequency of EM wave is swept from f_1 to f_2 , the phase change at the remote end can be obtained by demodulating the transmitted EM wave synchronously. From the cyclic change of the phase, we can obtain ΔN .

The proof-of-principle test has been done whose schematic is shown in Fig.3(a).

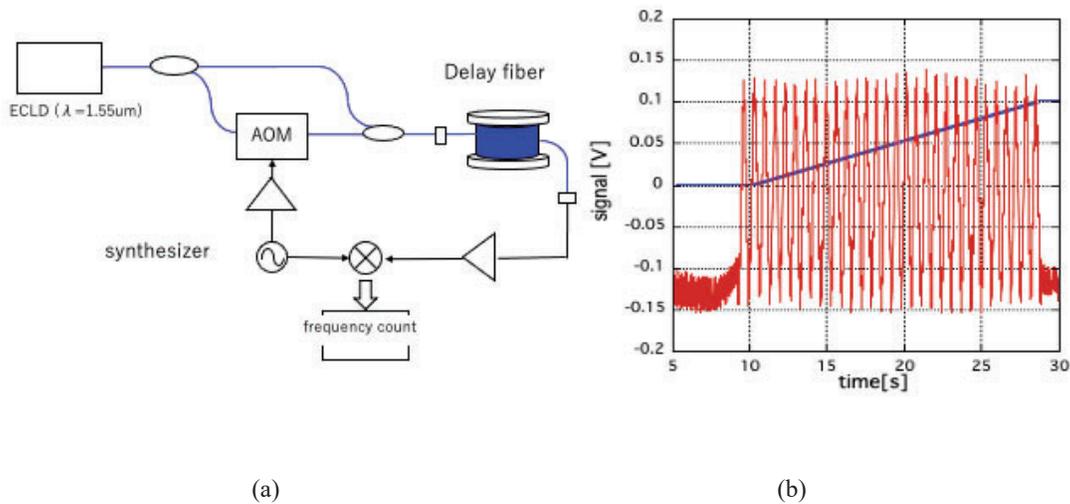


Figure 3. (a) Schematic diagram of the length measurement test , (b) temporal variation of the demodulated phase(red trace) and frequency of the beat signal (blue trace)

For long distance measurement, single-mode optical fibers are used as the optical path. 1.55- μm single-frequency light source from a fiber-delivered micro external-cavity laser diode is divided into two by a 10-dB optical coupler. A small portion of the light is frequency-shifted by using an in-line AOM, and is re-combined with the laser through another branch. The recombined beam which contains two frequency components is transmitted through a delivering optical fiber, and the transmitted light is detected by a photodetector at the end of the fiber to make beat note. The detected beat note is mixed with the local oscillator to obtain demodulated phase of the transmitted beat. The AOM driving frequency and the local oscillator is synchronously swept, and the temporal variation of the demodulated phase is recorded (red trace in Fig.3(b)). By using an up- and down-counter, we can measure N from the variation of the phase with the resolution of $N/2$, and the optical length of the fiber is calculated from the equation (1). We measure a various kind of fiber length by using combinations of 1-km, 4-km and 25-km fiber spools. The driving frequency is swept from 50 MHz to 66 MHz, and the length of the fibers are measured with length resolution of 6.25 m because no miscounting is occurred during measurement.

This length measurement method is applied to our angle measurement system as follows; after diffracted beam from the leader satellite hits the follower satellite, the AOD driving frequency is fixed to keep illuminating the follower satellite, and the frequency of the local oscillator is scanned by AOM1. The length resolution of this system can be improved by increasing the scanned frequency range Δf . In the case of our system, the maximum frequency scanning range is 100 MHz which corresponds to the length resolution of 1.5 m. The length resolution would be further improved by detecting residual phase in addition to N counting.

4. CONCLUSIONS

We have developed 3-dimensional satellite pointing system, in which the 2-dimensional angle and the distance between satellites can be measured. In the preliminary test, the angle precision of 0.003 mrad and the length resolution of 6.25 m are achieved. Currently the breadboard model of this system has been developed which is installed on aluminum breadboard and plan to test at the long vacuum tube. This system has been developed for initial alignment of DECIGO project, and would be a key technology for large-size precision formation flying missions.

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